

SCIDAC Center for Extended MHD Modeling

# EXTENDED MHD MODELING: BACKGROUND, STATUS AND VISION

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## **OVERVIEW**

- The Extended MHD model
- The computational challenges:
  - Extreme separation of time scales
  - Extreme separation of spatial scales
  - Extreme anisotropy
  - Importance of geometry, boundary conditions
  - Causality: can't parallelize over time!

At least as challenging as hydrodynamic turbulence!

- Present computational approaches:
  - Implicit time differencing
  - Specialized spatial grids
- Status of present models
- Vision for integrated modeling



## DEFINITIONS

- Hydrodynamics A mathematical model that describes the motion of a continuous, isotropic fluid
- Magnetohydrodynamics (MHD) A mathematical model that describes the motion of a continuous, electrically conducting fluid in a magnetic field
  - Hydrodynamics and Maxwell equations coupled through Lorentz body force and Ohm's law
- Ideal MHD the fluid has infinite electrical conductivity (zero resistivity)
- Resistive MHD The fluid has finite conductivity and resistivity
- Extended MHD additional effects of electron dynamics and/or non-Maxwellian species



### **MODERN TOKAMAKS ARE RICH IN MHD ACTIVITY**

Example: DIII-D shot 86144 18-Aug-00 09:04:47 86144 \*104 1.0 PINJ 3/2 NTM 2/1 wall locking 6.0 BETA(%) 4\*ht/(aB \*10-4 7.5 n=2 BRMS (T Mm \*10<sup>-3</sup> 1.5 n=1 BRMS (T) 1/1 sawteeth \*10<sup>16</sup> M 2.0 DALPHA Disruption 2000 1000 1500 2500 3000 3500 4000 TIME (msec)

- •Sawtoothing discharge •3/2 NTM triggered at 2250 msec
- •2/1 locks to the wall



### **MODELING REQUIREMENTS**

Slow evolution

Nonlinear fluid model required

• Plasma shaping

**Realistic geometry required** 

High temperature

Large "Reynolds' numbers"

Low collisionality

**Extensions to resistive MHD required** 

Strong magnetic field

Highly anisotropic transport required

Resistive wall

Non-ideal boundary conditions required



### **APPROACHES**

• Quasi-equilibrium:

 $\nabla p = \mathbf{J} \times \mathbf{B} + \text{constraints}$ 

- "Magnetohydrostatics" (MHS?)
- Eliminates all waves
- Basis for 1-1/2 dimensional transport models
- Extension to 3-D?
- Time dependent
  - Solve 2-fluid equations
  - Retain all normal modes
  - Focus of present SciDAC efforts



### **FLUID MODELS**

- Kinetic models of plasmas based on distribution function for • each charge species
- Satisfies kinetic equation ۲

$$\frac{df_{\alpha}}{dt} = \sum_{\beta} C[f_{\alpha}, f_{\beta}]$$

 $f_{\alpha}(\mathbf{x}, \mathbf{v}, t)$  - six dimensions plus time - computationally impractical for time scales of interest

- Fluid models derived by taking successive velocity moments of kinetic equation
  - Reduce dimensionality by 3
- Hierarchy of equations for  $n, v, p, \Pi, q, \dots$
- Equations truncated by closure relations
  - Express high order moments in terms of low order moments
  - Capture *kinetic effects* in these moments
- Result is Extended MHD



• Maxwell (no displacement current):

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}$$
 ,  $\nabla \times \mathbf{B} = \mu_0 \mathbf{J}$ 

• Momentum, energy, and continuity for each species ( $\alpha = e, i$ ):

$$m_{\alpha}n_{\alpha}\left(\frac{\partial \mathbf{v}_{\alpha}}{\partial t} + \mathbf{v}_{\alpha} \cdot \nabla \mathbf{v}_{\alpha}\right) = -\nabla \cdot \mathbf{P}_{\alpha} + q_{\alpha}n_{\alpha}\left(\mathbf{E} + \mathbf{v}_{\alpha} \times \mathbf{B}\right) + \sum_{\beta}\mathbf{R}_{\alpha\beta} + \mathbf{S}_{\alpha}^{m}$$
$$\frac{\partial p_{\alpha}}{\partial t} + \mathbf{v}_{\alpha} \cdot \nabla p_{\alpha} = -\frac{3}{2}p_{\alpha}\nabla \cdot \mathbf{v}_{\alpha} - \mathbf{P}_{\alpha} : \nabla \mathbf{v}_{\alpha} - \nabla \cdot \mathbf{q}_{\alpha} + \mathbf{Q}_{\alpha}$$
$$\frac{\partial n_{\alpha}}{\partial t} = -\nabla \cdot (n_{\alpha}\mathbf{v}_{\alpha}) + \mathbf{S}_{\alpha}^{n}$$

• Current and quasi-neutrality:

$$\mathbf{J}_{\alpha} = n_{\alpha} q_{\alpha} \mathbf{v}_{\alpha}, \qquad n = n_{\mathbf{e}} = Z n_{j}$$



• Add electron and ion momentum equations:

$$\rho \left( \frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} \right) = -\nabla \cdot \mathbf{P}' + \mathbf{J} \times \mathbf{B}$$

• Subtract electron and ion momentum equations (Ohm's law):



### All effects beyond resistivity constitute Extended MHD



### **COMPUTATIONAL CHALLENGES**

- Extreme separation of time scales
  - Realistic "Reynolds' numbers"
  - Implicit methods
- Extreme separation of spatial scales
  - Important physics occurs in internal boundary layers
  - Small dissipation cannot be ignored
  - Requires grid packing or adaption
- Extreme anisotropy
  - Special direction determined by magnetic field Accurate treatment of  $\mathbf{B} \cdot \nabla$  operator is important
  - Requires specialized gridding



### **SEPARATION OF TIME SCALES**



Lundquist number: 
$$S = \frac{\tau_R}{\tau_A} \sim 10^8 >> 1$$

**Explicit time step impractical:** 

$$\Delta t < \frac{\Delta x}{L} \tau_A \approx \frac{\tau_A}{N} <<<< au_{evol}$$

#### **Require implicit methods**



### **IMPLICIT METHODS**

- Partially implicit methods
  - Treat fastest time scales implicitly
  - Time step still limited by waves
- Semi-implicit methods
  - Treat linearized ideal MHD operator implicitly
  - Time step limited by advection
  - Many iterations
- Fully implicit methods
  - Newton-Krylov treatment of full nonlinear equations
  - Arbitrary time step
  - Still a research project



### LINEAR SOLVER REQUIREMENTS

- Extremely large condition number : > 10<sup>10</sup>!!
  - Specialized pre-conditioners
  - Anisotropy
- Ideal MHD is self-adjoint
  - Symmetric matrices
  - CG
- Advection and some 2-fluid effects (whistler waves) are not selfadjoint
  - Need for efficient non-symmetric solvers
- Everything must be efficient and scalable in parallel
- Should interface easily with F90



### **SEPARATION OF SPATIAL SCALES**

- Important dynamics occurs in internal boundary layers
  - Structure is determined by plasma resistivity or other dissipation
  - Small dissipation cannot be ignored
- Long wavelength along magnetic field
- Extremely localized across magnetic field:

$$\delta IL \sim S^{-\alpha} \ll 1 \text{ for } S \gg 1$$

 It is these long, thin structures that evolve nonlinearly on the slow evolutionary time scale



### **EXTREME ANISOTROPY**

- Magnetic field locally defines special direction in space
- Important dynamics are extended along field direction, very narrow across it
- Propagation of normal modes (waves) depends strongly on local field direction
- Transport (heat and momentum flux) is also highly anisotropic

==> Requires accurate treatment of operator  $\mathbf{B} \cdot \nabla$ 

Inaccuracies lead to "spectral pollution" and anomalous perpendicular transport



### **GRIDDING AND SPATIAL REPRESENTATION**

- Spatial stiffness and anisotropy require special gridding
  - Toroidal and poloidal dimensions treated differently
- Toroidal ( $\phi$ , primarily along field)
  - Long wavelengths, periodicity => FFTs (finite differences also used)
- Poloidal plane (*R*,*Z*)
  - Fine structure across field direction
  - Grids aligned with flux surfaces (~ field lines)
  - Unstructured triangular grids
  - Extreme packing near internal boundary layers
- Finite elements
  - High order elements essential for resolving anisotropies
- Dynamic mesh adaption in research phase



### **POLOIDAL GRIDS**





DIII-D poloidal cross-section with flux aligned grid (NIMROD)

Circular poloidal cross-section with triangles and grid packing (M3D)

### Poloidal grids from SciDAC development projects



### **BEYOND RESISTIVITY - EXTENDED MHD**

- 2-fluid effects
  - Whistler waves (Hall term) require implicit advance with non-symmetric solver
  - Electron inertia treated implicitly
  - Diamagnetic rotation may cause accuracy, stability problems
- Kinetic effects influence of non-Maxwellian populations
  - Analytic closures
    - Seek *local* expressions for  $\Pi$ , q, etc.
  - Particle closures
    - Subcycle gyrokinetic  $\delta f$  calculation
    - Minority ion species beam or  $\alpha$ -particles



## **STATUS**

- 2 major SciDAC development projects for time-dependent models
  - M3D multi-level, 3-D, parallel plasma simulation code
    - Partially implicit
    - Toroidal geometry suitable for stellarators
    - 2-fluid model
    - Neo-classical and particle closures
  - NIMROD 3-D nonlinear extended MHD
    - Semi-implicit
    - Slab, cylindrical, or axisymmetric toroidal geometry
    - 2-fluid model (evolving computationally)
    - Neo-classical closures
    - Particle closures being de-bugged

Both codes have exhibited good parallel performance scaling

• Other algorithms are being developed in the fusion program



### **STATUS - RESISTIVE MHD**





Sawtooth in NSTX computed by M3D

#### Stellarator ballooning mode computed by M3D



### **STATUS - RESISTIVE MHD**





Secondary magnetic islands generated during sawtooth crash in DIII-D shot 86144 by NIMROD



### **STATUS - EXTENDED MHD**

- Effect of energetic particle population on MHD mode
- Subcycling of energetic particle module within MHD codes
- M3D agrees well with NOVA2 in the linear regime
- Energetic particles are being incorporated into NIMROD





### **STATUS - EXTENDED MHD**

### Neo-classical tearing modes with NIMROD using analytic closure





### **NEXT STEP - INTEGATED MODELING**

- Non-local kinetic physics, MHD, and profile evolution are all interrelated
  - Kinetic physics determines transport coefficients
  - Transport coefficients affect profile evolution
  - Profile evolution can destabilize of MHD modes
  - Kinetic physics can affect nonlinear MHD evolution (NTMs, TAEs)
  - MHD relaxation affects profile evolution
  - Profiles affect kinetic physics
- Effects of kinetic (sub grid scale) physics must be synthesized into MHD models
  - Extensions to Ohm's law (2-fluid models)
  - Subcycling/code coupling
  - Theoretical models (closures), possibly heuristic
- Effects of MHD must be synthesized into transport models
- Predictions must be validated with experimental data



### **VISION: VDE EVOLUTION**





### **VISION: SAWTOOTH CYCLE**





### **ENABLING COMPUTER SCIENCE TECHNOLGIES**

- Largest, fastest computers!
  - But intermediate computational resources often neglected, and...
  - The computers will never be large or fast enough!
- Algorithms
  - Parallel linear algebra
  - Gridding, adaptive and otherwise
- Data structure and storage
  - Adequate storage devices
  - Common treatment of experimental and simulation data
  - Common tools for data analysis
- Communication and networking
  - Fast data transfer between simulation site and storage site
  - Efficient worldwide access to data
  - Collaborative tools
  - Dealing with firewalls
- Advanced graphics and animation



### SUMMARY

- Predictive simulation capability has 3 components
  - Code and algorithm development
  - Tightly coupled theoretical effort
  - Validation of models by comparison with experiment
- Integration required for:
  - Coupling algorithms for disparate physical problems
  - Theoretical synthesis of results from different models
  - Efficient communication and data manipulation
- Progress is being made with Extended MHD
  - Integration of energetic ion modules into 3-D MHD
  - Computationally tractable closures
  - Resistive wall modules

Need to bring a broader range of algorithms and codes to bear for overall fusion problem

